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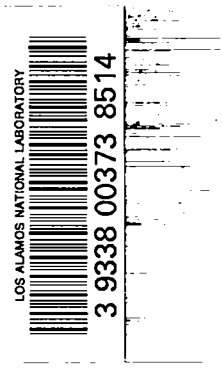
Thermal Neutron Spectra from an Underground Nuclear Explosion with Special Consideration of Spectral Modification due to Bomb Debris Motion



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Thermal Neutron Spectra from an Underground Nuclear Explosion with Special Consideration of Spectral Modification due to Bomb Debris Motion

by

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ABSTRACT

The neutron flux from an underground nuclear explosion has recently been used to measure nuclear cross sections. For the experimenter it is of interest to know the neutron spectrum and intensity emerging from the top of the pipe leading to the nuclear device.

A short time following the nuclear explosion and after many fast neutrons have escaped, the local neutron spectra in the bomb debris will be in thermal equilibria. The spectrum of neutrons leaking from this debris will depend on a superposition of these local Maxwell Boltzmann distributions each with its characteristic drift velocity $\vec{v}_o$.

Theoretical expressions are derived for the low energy neutron spectrum emerging from a long narrow pipe. The theoretical predictions agree reasonably well with measurements obtained from the Parrot Shot in Nevada in the Spring of 1965.

ACKNOWLEDGMENT

We would like to thank our colleagues, Walter Goad and George Bell, for their help in this investigation.

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I. INTRODUCTION

The problem to be solved is: given a spherical ball of bomb debris moving radially outwards with material velocity $\vec{v}_o$ and containing neutrons in local thermal equilibria, then what is the neutron spectrum on top of a long narrow pipe of length R and radius η (Fig. 1). It is assumed that the debris density has become so small that no further neutron collisions will occur; in other words, the neutrons do not find a diffusive medium anymore. The neutrons are therefore liberated from a moving source containing "bottled up" neutrons with local Maxwell Boltzmann distributions. The spectrum on top of the pipe will depend on how much one "sees" of the bomb debris sphere looking down the long narrow pipe. This problem has been discussed previously by Nordheim,[1] as well as Bell.[2] This paper looks at the problem in a more general way, and the case treated by Nordheim and Bell becomes here a special case where the observer looking down the pipe sees only a "pencil" of the bomb debris sphere.

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[1] L. W. Nordheim: Neutron Spectra from a Moving Source, Los Alamos Scientific Laboratory Memorandum, July 30, 1964, and Addendum, July 22, 1964.

[2] G. I. Bell, private communication, August 19, 1964.

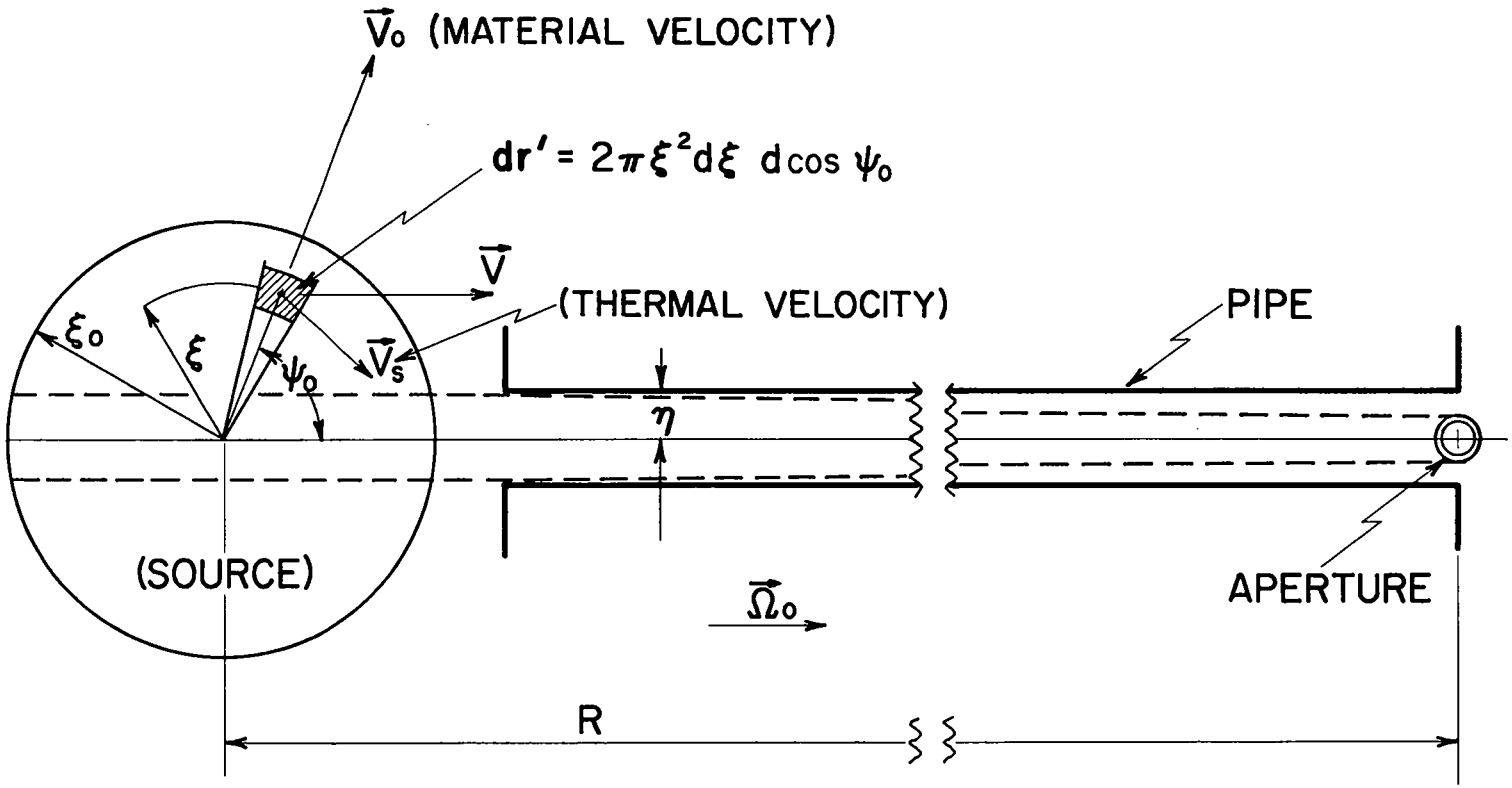

FIG. 1. SCHEMATIC OF DEBRIS SPHERE AND PIPE.

In Section II we derive the general case where the observer sees a general portion of the debris sphere.

Subsequently in Sections III and IV we discuss two special cases called the "pencil" model and the "spherical" model. As the name implies, we will see the whole sphere in the "spherical" model, and in the "pencil" model we only see a small pencil-like section of the debris sphere. In Section V we give the numerical bomb parameters relevant to this investigation. The theoretical predictions of the neutron spectrum on top of the pipe are compared with the experimental data from the "Parrot" shot.

The effect of the finite bomb debris sphere on measurements is investigated in Section VI.

II. THERMAL NEUTRON SPECTRA FROM A RAPIDLY EXPANDING MATERIAL SPHERE

The initial neutron density in the debris is assumed to be a Maxwell-Boltzmann distribution.

$$n(\vec{v}_s) = ce^{-\frac{1}{2}m\frac{(\vec{v}-\vec{v}_o)^2}{KT}} = ce^{-\frac{1}{2}\frac{m\vec{v}_s^{\,2}}{KT}} \qquad \text{EQN 1}$$

Here the drift velocity is equal to the material velocity $\vec{v}_o$, and $\vec{v}_s$ is the neutron velocity relative to $\vec{v}_o$. In Appendix A we derived an expression for the number of neutrons passing through the unit area perpendicular to direction $\vec{\Omega}_o$ at R (Fig. 1) in the time interval ΔT about T. The initial neutron population $n(\vec{r}',v,\vec{\Omega}_o,0^+)$ in Eqn. A-14 created by the δ-function source becomes

$$n(\vec{r}',v,\vec{\Omega}_o,0^+) = c(\vec{\xi})e^{-\frac{1}{2}\frac{m}{KT}[v^2+v_o^{\,2}-2vv_o\cos\psi_o]} \times v^2 \qquad \text{EQN 2}$$

and Eqn. A-14 becomes

$$J_o(\vec{r},T)\Delta T = \int_{\text{source}} 2\pi\xi^2 d\xi\, d\cos\psi_o\, c(\vec{\xi})e^{-\frac{m}{2KT}[v^2+v_o^{\,2}-2vv_o\cos\psi_o]} \times v^2 \frac{\Delta v}{R^2} \qquad \text{EQN 3}$$

CASE I: uniform gas, i.e. $c(\vec{\xi}) = \text{const.}$

from $\xi = 0$ to $\xi = \xi_o$; $v_o = \text{const.}$

The assumption made above means that the neutron production is uniform over the debris sphere and the debris velocity is constant in the radial direction. The evaluation of Eqn. 3 is split into three intervals (Fig. 2), i.e. regions I, II, and III. The limits of variable $\cos\psi_o$ in the integrals are reversed because $-d\cos\psi_o = \sin\psi_o d\psi_o$. We get in the three regions

$$\frac{J_o(\vec{r},T)\Delta T}{\Delta v} = 2\pi \int_o^{\xi_o} \xi^2 d\xi \, \frac{cv^2}{R^2} \left[\int_{\cos\psi_{oo}}^{1.0} d\cos\psi_o e^{-\frac{m}{2KT}[v^2+v_o^2-2vv_o\cos\psi_o]} + \right.$$

$$\left. + \int_{-1.0}^{-\cos\psi_{oo}} d\cos\psi_o e^{-\frac{m}{2KT}[v^2+v_o^2-2vv_o\cos\psi_o]} \right] +$$

$$+ 2\pi \int_{-\cos\psi_{oo}}^{+\cos\psi_{oo}} d\cos\psi_o \, \frac{cv^2}{R^2} e^{-\frac{m}{2KT}[v^2+v_o^2-2vv_o\cos\psi_o]} \int_o^{\eta/\sin\psi_o} \xi^2 d\xi$$

EQN 4

In the 3rd term in Eqn. 4 we assume that

$$\int_o^{\eta/\sin\psi_o} \xi^2 d\xi = \frac{\eta^3}{\langle \sin^3\psi_o \rangle 3}$$

EQN 5

and that the average value of $\sin\psi_o$ or $\langle \sin\psi_o \rangle$ can be found graphically over region III in Fig. 2. This step is convenient to obtain a

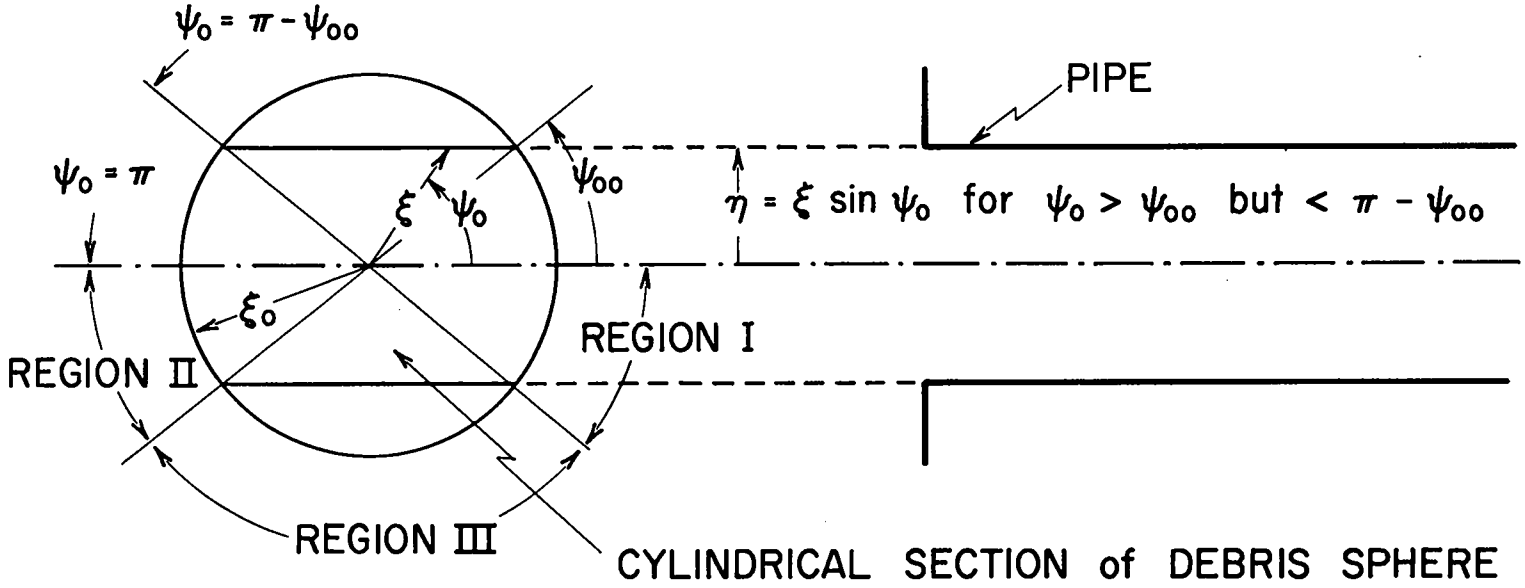

FIG. 2. CYLINDRICAL SECTION OF DEBRIS SPHERE AND PIPE.

simple solution for the 3rd term in Eqn. 4.

After solving for the definite integrals in Eqn. 4 we get

$$\frac{J_o(\vec{r},T)\Delta T}{\Delta v} = \frac{2\pi\xi_o^3 cvKT}{3m\, v_o\, R^2}\left[e^{-\frac{m}{2KT}(v-v_o)^2} - e^{-\frac{m}{2KT}(v^2+v_o^2-2vv_o\cos\psi_{oo})} + e^{-\frac{m}{2kt}(v^2+v_o^2+2vv_o\cos\psi_{oo})} - e^{-\frac{m}{2KT}(v+v_o)^2}\right] + \frac{2\pi\eta^3 cvKT}{3mv_o\langle\sin^3\psi_o\rangle R^2}\left[e^{-\frac{m}{2KT}(v^2+v_o^2-2vv_o\cos\psi_{oo})} - e^{-\frac{m}{2KT}(v^2+v_o^2+2vv_o\cos\psi_{oo})}\right] \qquad \text{EQN 6}$$

or rearranging

$$\frac{J_o(\vec{r},T)\Delta T}{\Delta v} = \frac{2\pi\xi_o^3 cvKT}{3mv_o R^2}\left\{e^{-\frac{m}{2KT}(v-v_o)^2} - e^{-\frac{m}{2KT}(v+v_o)^2} + \left[\frac{\eta^3}{\xi_o^3\langle\sin^3\psi_o\rangle} - 1\right]\left[e^{-\frac{m}{2KT}(v^2+v_o^2-2vv_o\cos\psi_{oo})} - e^{-\frac{m}{2KT}(v^2+v_o^2+2vv_o\cos\psi_{oo})}\right]\right\} \qquad \text{EQN 7}$$

From Appendix B, Eqn. B-10, we get for the constant c in Eqn. 7

$$c = \frac{N_o}{V_o} \left(\frac{2\pi KT}{m}\right)^{-3/2} \qquad \text{EQN 8}$$

where

N_o = total thermal neutrons produced in sphere

V_o = sphere volume in cm^3

We insert Eqn. 8 into Eqn. 7

$$\frac{J_o(\vec{r},T)\Delta T}{\Delta v} = \frac{N_o}{4\pi R^2} \frac{v}{v_o} \left(\frac{m}{2\pi KT}\right)^{1/2} \left\{ e^{-\frac{m}{2KT}(v-v_o)^2} - e^{-\frac{m}{2KT}(v+v_o)^2} + \right.$$

$$+ \left[\frac{\eta^3}{\xi_o^3 \langle \sin^3 \psi_o \rangle} - 1\right] \left[e^{-\frac{m}{2KT}(v^2+v_o^2-2vv_o\cos\psi_{oo})} \right.$$

$$\left. \left. - e^{-\frac{m}{2KT}(v^2+v_o^2+2vv_o\cos\psi_{oo})} \right] \right\} \qquad \text{EQN 9}$$

We introduce the following relationship between v_o, the material debris velocity, and the neutron temperature KT

$$\frac{mv_o^2}{2KT} = \beta \qquad \text{EQN 10}$$

also

$$\upsilon = \frac{v}{v_o} \qquad \text{EQN 11}$$

We insert Eqns.10 and 11 into Eqn. 9 and get finally for the general

case

$$J_o(\vec{r},T)\Delta T = \frac{N_o}{4\pi R^2}\sqrt{\frac{\beta}{\pi}}\,\upsilon\Delta\upsilon \left\{ e^{-\beta(\upsilon-1)^2} - e^{-\beta(\upsilon+1)^2} + \right.$$

$$+ \left[\frac{\eta^3}{\xi_o^3\langle \sin^3\psi_o\rangle} - 1\right]\left[e^{-\beta(\upsilon^2+1-2\upsilon\cos\psi_{oo})}\right.$$

$$\left.\left. - e^{-\beta(\upsilon^2+1+2\upsilon\cos\psi_{oo})}\right]\right\} \qquad \text{EQN 12}$$

Equation 12 gives the number of neutrons passing through the unit area perpendicular to $\vec{\Omega}_o$ at $\vec{r}$ (Fig. A-1) in the time interval ΔT about T in terms of that portion of the initial neutron population at $\vec{r}'$ with energies specified by $\Delta\upsilon$ about υ. This is for the general case in Fig. 2 where an observer, looking down the pipe, "sees" a general portion of the neutron generating sphere.

This general case in Eqn. 12 is not exploited numerically in this paper, but in the following we treat two practical cases which represent the two limiting cases to the general case expressed in Eqn. 12. The first case is called the "pencil" model where the observer looking down the pipe sees only a small narrow pencil-like section of the neutron producing sphere. The second case is called the "spherical" model where the observer sees the whole neutron producing sphere.

III. "PENCIL" MODEL

Here the observer sees only a small narrow pencil-like section of the neutron producing sphere (Fig. 3). For this case it follows from Eqn. 3, Eqn. B-10, and evaluating the exponential term in Eqn. 3 for $\cos\psi_o \sim 1$ and $\cos\psi_o \sim -1$

$$J_o(\vec{r},T)\Delta T = \frac{N_o}{V_o}\left(\frac{2\pi KT}{m}\right)^{-3/2}\frac{U_o v^2 \Delta v}{2R^2}\left[e^{-\frac{m}{2KT}(v-v_o)^2} + e^{-\frac{m}{2KT}(v+v_o)^2}\right]$$

EQN 13

Multiplying Eqn. 13 by πr^2, the aperture area of the pipe, and using the bomb parameter value $\beta = mv_o^2/2KT = 1$ (see Section V), we have

$$\pi r^2 J_o(\vec{r},T)\Delta t = \frac{r^2}{4R^2}\frac{N_o}{\sqrt{2\pi}}2\sqrt{2}\frac{U_o}{V_o}\upsilon^2\Delta\upsilon\left[e^{-(\upsilon-1)^2} + e^{-(\upsilon+1)^2}\right] \qquad \text{EQN 14}$$

where U_o = pencil volume in cm^3, V_o = sphere volume in cm^3, N_o = total thermal neutrons produced, $\upsilon = v/v_o$, and $v_o = \sqrt{2KT/m}$. Equation 14 gives the total neutrons passing through the aperture area πr^2 (cm^2) on top of the pipe perpendicular to $\vec{\Omega}_o$ at distance R (cm) from the source.

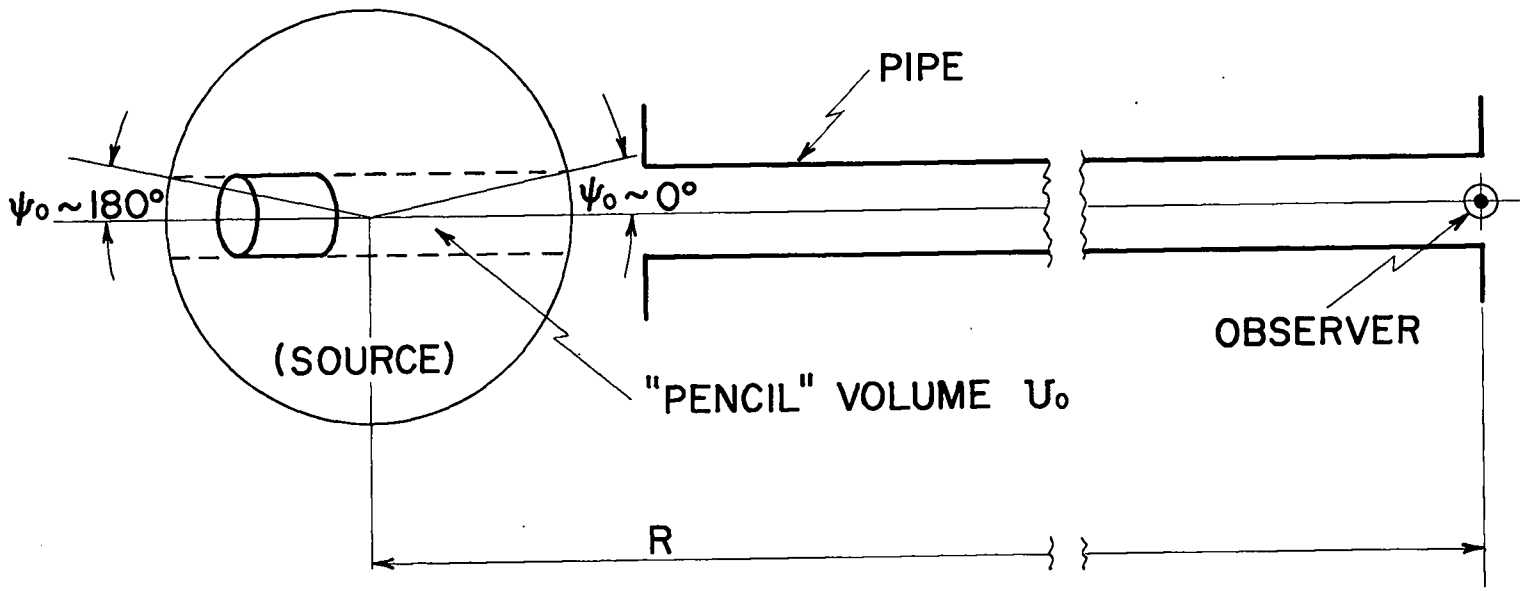

FIG. 3. "PENCIL" MODEL OF DEBRIS SPHERE AND PIPE.

For comparison, we give below the corresponding equation for the hypothetical case of $v_o = 0$.

$$\pi r^2 J_o(\vec{r},T)\Delta T = \frac{r^2}{2R^2} \frac{N_o}{\sqrt{2\pi}} \frac{1}{\sqrt{\frac{KT}{m}}\frac{KT}{m}} \frac{U_o}{V_o} v^2 \Delta v e^{-\frac{mv^2}{2KT}} \qquad \text{EQN 14a}$$

IV. "SPHERE" MODEL

Here the observer sees the whole neutron producing sphere (Fig. 4). For this case it follows from Eqn. 3, Eqn. B-10 and integrating $\cos\psi_o$ in Eqn. 3 over the whole range + 1 to - 1

$$J_o(\vec{r},T)\Delta T = N_o\left(\frac{2\pi KT}{m}\right)^{-3/2} \frac{v^2\Delta v}{2R^2}\left\{\frac{KT}{mv_o v}\left[e^{-\frac{m}{2KT}(v-v_o)^2} - e^{-\frac{m}{2KT}(v+v_o)^2}\right]\right\} \qquad \text{EQN 15}$$

As in the pencil model we assume again that $\beta = 1$, and, after multiplying Eqn. 15 by πr^2, obtain

$$\pi r^2 J_o(\vec{r},T)\Delta T = \frac{r^2}{4R^2}\,\frac{N_o}{\sqrt{\pi}}\,\upsilon^2\Delta\upsilon\left\{\frac{1}{\upsilon}\left[e^{-(\upsilon-1)^2} - e^{-(\upsilon+1)^2}\right]\right\} \qquad \text{EQN 16}$$

where $\upsilon = \frac{v}{v_o}$, $v_o = \sqrt{\frac{2KT}{m}}$, N_o = total thermal neutrons produced.

Equation 16 gives the total neutrons passing through the aperture area πr^2 (cm^2) on top of the pipe perpendicular to $\vec{\Omega}_o$ at distance R (cm) from the source.

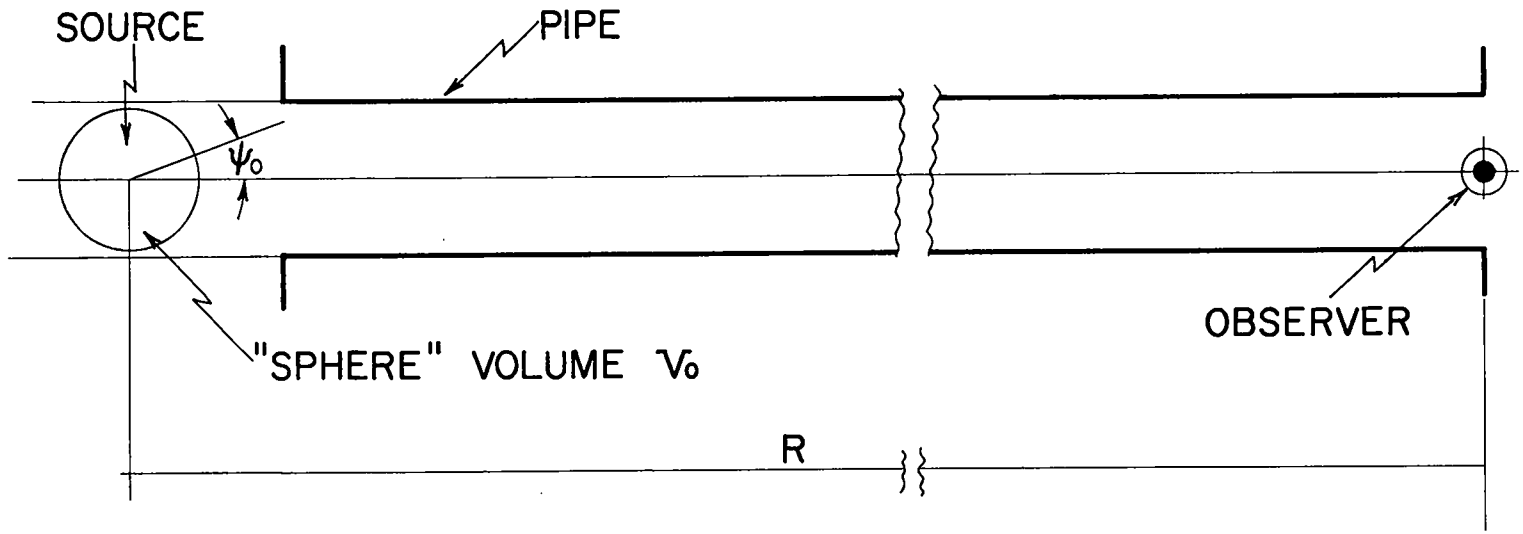

FIG. 4. "SPHERE" MODEL OF DEBRIS SPHERE AND PIPE.

For comparison, we again give below the corresponding equation for the hypothetical case of $v_o = 0$.

$$\pi r^2 J_o(\vec{r},T)\Delta T = \frac{r^2}{2R^2} \frac{N_o}{\sqrt{2\pi}} \frac{v^2 \Delta v}{\sqrt{\frac{KT}{m}} \frac{KT}{m}} e^{-\frac{m}{2KT} v^2} \qquad \text{EQN 16a}$$

V. NUMERICAL EVALUATION OF "PARROT" ESCAPE SPECTRUM

The experimental set-up (Fig. 5) has been described in the open literature.[3]

The numerical values for the quantities shown in Fig. 5 are listed below. These values are subsequently used to numerically evaluate Eqn. 14 "pencil" model and Eqn. 16 "sphere" model.

$N_o = 0.06848 \times 10^{24}$	"Parrot" neutrons released in thermal equilibrium
$V_o = 5.24 \times 10^5$ cm^3	Bomb debris "sphere" volume at time of neutron release
$U_o = 3.24 \times 10^4$ cm^3	Bomb debris "pencil" volume at time of neutron release
R = 182.18 m	Distance bomb debris center to fission foil detector of area $\pi r^2 = 1.037$ cm^2 pipe radius $\eta = 10.16$ cm
Y = 1.2 KT	Total yield of "Parrot"
KT = 474 ev	Bomb thermal temperature
$v_o = 0.3$ cm/sh	Material velocity

These individual quantities shown in Fig. 5 will now be discussed briefly.

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[3] "Neutron Time of Flight Experiments Using Nuclear Explosions," B. C. Diven. Submitted and to be published in Oct. 1965 issue of SCIENCE.

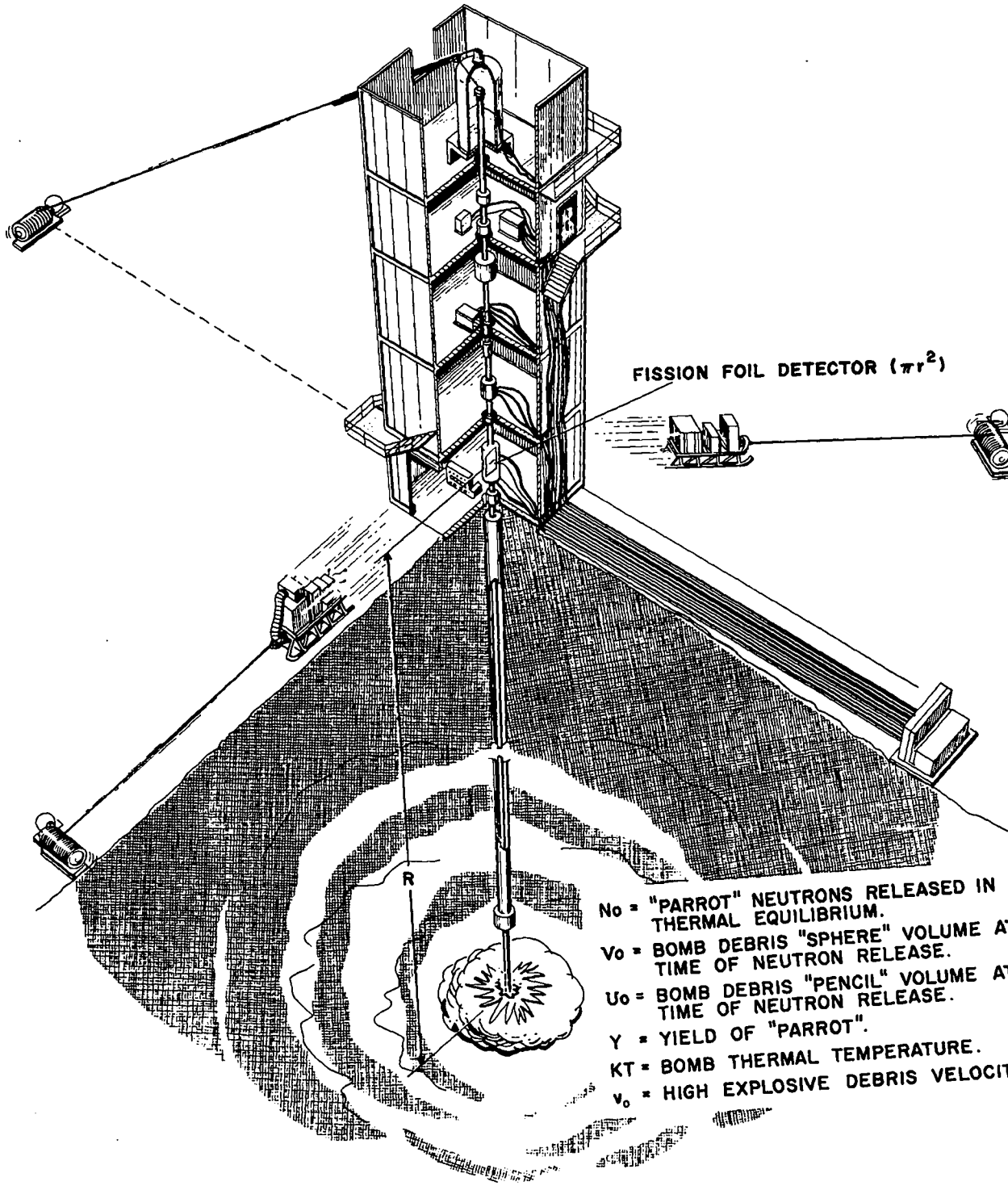

FIG. 5. UNDERGROUND NUCLEAR TEST "PARROT."

a. N_o, Parrot neutrons released in thermal equilibrium; v_o, debris velocity and KT, Parrot bomb thermal temperature.

In a nuclear device a configuration of fissionable material is driven together by exploding the surrounding high explosive layer.

The assembled configuration will reach supercriticality and the released fission energy will cause the rapid expansion of the bomb. The debris will move radially outward, including the hydrogeneous remnants of the high explosive layer. Before the high explosive moderator or reflector moves away, it is assumed that the fission neutrons born have come to a thermal equilibrium by collisions with the bomb debris.

During the expansion phase a time will be reached where the surrounding moderator (high explosive remnants) comes to such a low density that the thermal neutrons will cease to make collisions with the high explosive debris. We are therefore confronted with the problem to determine the neutron spectrum coming from a moving source, as seen in the Laboratory system. The material velocity v_o of the debris becomes the source velocity of the "bottled up" thermal neutrons.

From coupled hydrodynamics--neutronics calculations--the number of thermal neutrons N_o can be determined as well as the velocity v_o of the high explosive debris. A reasonable estimate of the HE debris velocity v_o at the time of thermal neutron release is 0.3 cm/shake. The debris temperature may now be deduced by using the fact that, behind a strong shock, internal energy per gram equals kinetic energy per gram: for the

essentially completely ionized HE components, this translates to $v_o \simeq \sqrt{2KT/m}$ with m still representing the mass of one neutron. We thus obtain $KT \simeq 474$ ev.

b. Pencil and sphere volume, V_o and U_o.

The volume of the debris sphere U_o, as well as the pencil volume V_o, is determined from the same calculations mentioned previously. The sphere volume V_o is the configuration in which the mean free path for thermal neutrons in the high explosive remnants becomes so large that it will no longer act as a diffusive medium for thermal neutrons.

c. Depth of underground nuclear explosion, R.

The quantity of interest in this report is R = 182.18 m and the foil area $\pi r^2 = 1.037$ cm^2. The fixed aperture area just below ground level is 0.9698 cm^2. The fission foil is a few feet above ground level.

d. Evaluation of "pencil" and "sphere" neutron spectrum.

Equations 14 and 14a, and Eqns. 16 and 16a are evaluated for even steps in the dimensionless variable $\upsilon = v/v_o$ (Table 1). For $\upsilon = 1$ the neutron velocity v traveling up the pipe is equal to v_o, the debris velocity. Table 1 assumes a bomb thermal temperature of 474 ev. Columns 2 through 8 are evaluated from the following relationships

$$v = \upsilon \times v_o = 0.3\upsilon \text{ (cm/sh)}$$

$$\Delta v = \Delta\upsilon \times v_o = 0.3\Delta\upsilon \text{ (cm/sh)}$$

$$E(\text{Mev}) = \frac{[v(\text{cm/sh})]^2}{(13.827)^2}$$

$$\Delta E(\text{Mev}) = v(\text{cm/sh})\ \Delta v(\text{cm/sh})\ \frac{2}{(13.827)^2}$$

$$T(\text{sh}) = \frac{R(\text{cm})}{v(\text{cm/sh})} = \frac{1.8218 \times 10^4}{v(\text{cm/sh})}$$

$$\Delta T(\text{sh}) = \frac{R(\text{cm})\Delta v(\text{cm/sh})}{[v(\text{cm/sh})]^2} = \frac{1.8218 \times 10^4\ \Delta v(\text{cm/sh})}{[v(\text{cm/sh})]^2}$$

Note that the intervals Δv, ΔE and ΔT follow from the arbitrarily chosen interval $\Delta \upsilon$.

The numerical expression to evaluate $J_o(\vec{r},T)$ in Table 1 for all models and assumptions are listed below.

"Pencil" model; $v_o = 0.3$ cm/sh; $v_o = \sqrt{\frac{2KT}{m}}$

Table 1, Column 9. $KT = 474$ ev.

From Eqn. 14 we get

$$\pi r^2 J_o(\vec{r},T)\Delta T = 1.1878 \times 10^{12}\ \upsilon^2 \Delta\upsilon \left[e^{-(\upsilon-1)^2} + e^{-(\upsilon+1)^2} \right] \qquad \text{EQN 17}$$

- -

"Pencil" model; $v_o = 0$, Maxwellian; $\sqrt{\frac{2KT}{m}} = 0.3$ cm/sh

Table 1, Column 10. $KT = 474$ ev.

From Eqn. 14a we get

$$\pi r^2 J_o(\vec{r},T)\Delta T = 8.7985 \times 10^{-11}\ v^2 \Delta v e^{-1.1111 \times 10^{-15}\ v^2} \qquad \text{EQN 18}$$

"Sphere" model; $v_o = 0.3$ cm/sh; $v_o = \sqrt{\frac{2KT}{m}}$

Table 1, Column 11. $KT = 474$ ev.

From Eqn. 16 we get

$$\pi r^2 J_o(\vec{r},T)\Delta T = 9.6047\times10^{12}\, \upsilon^2 \Delta\upsilon \left\{\frac{1}{\upsilon}\left[e^{-(\upsilon-1)^2} - e^{-(\upsilon+1)^2}\right]\right\} \quad \text{EQN 19}$$

"Sphere" model; $v_o = 0$, Maxwellian; $\sqrt{\frac{2KT}{m}} = 0.3$ cm/sh

Table 1, Column 12. $KT = 474$ ev.

From Eqn. 16a we get

$$\pi r^2 J_o(\vec{r},T)\Delta T = 1.4263\times10^{-9}\, v^2 \Delta v\, e^{-1.1111\times10^{-15} v^2} \quad \text{EQN 20}$$

For our case $\pi r^2 = 1.037$ cm^2 and ΔT is taken from Table 1.

The last four columns of Table 1, as well as the experimental data, are plotted in Fig. 6.

The experimental data agree reasonably well with the calculated pencil model values. The neutron flux is "contaminated" with "reflected" neutrons which are not included in our calculations. The first peak at 18 ev is probably due to reflected neutrons from the bottom of the hole which picked up some energy from the "pre-heated" lining of the cavity. At higher energies above 3 kev we get the neutrons which have been slowed down from higher energies. This energy range is not included in this discussion, as we limited ourselves to the thermal part.

u	Δu	v (cm/sh)	Δv (cm/sh)	E	ΔE (ev)	T (msec)	ΔT (μsec)	$J_o(\vec{r},T)$ Neutrons/cm^2 sec; KT = 474 ev; R = 182.18 m			
								"Pencil"		"Sphere"	
								v_o=0.3 cm/sh	Maxwell v_o=0	v_o=0.3 cm/sh	Maxwell v_o=0
0.1	0.05	0.03	0.015	0.47 ev	4.71	6.073	3030	1.405×10^{11}	3.742×10^{11}	2.242×10^{12}	6.067×10^{12}
0.2	0.05	0.06	0.015	18.8 ev	9.41	3.036	759.1	2.307×10^{12}	5.799×10^{12}	3.544×10^{13}	9.400×10^{13}
0.3	0.05	0.09	0.015	42.4 ev	14.12	2.024	337.4	1.218×10^{13}	2.792×10^{13}	1.763×10^{14}	4.526×10^{14}
0.4	0.05	0.12	0.015	75.3 ev	18.83	1.518	189.8	4.048×10^{13}	8.228×10^{13}	5.432×10^{14}	1.334×10^{15}
0.5	0.05	0.15	0.015	117.7 ev	23.54	1.215	121.5	1.042×10^{14}	1.836×10^{14}	1.283×10^{15}	2.976×10^{15}
0.6	0.05	0.18	0.015	169.5 ev	28.24	1.012	84.3	2.273×10^{14}	3.413×10^{14}	2.554×10^{15}	5.533×10^{15}
0.7	0.05	0.21	0.015	230.6 ev	32.95	0.868	62.0	4.389×10^{14}	5.546×10^{14}	4.487×10^{15}	8.990×10^{15}
0.8	0.05	0.24	0.015	301.2 ev	37.66	0.759	47.4	7.732×10^{14}	8.155×10^{14}	7.203×10^{15}	1.322×10^{16}
0.9	0.05	0.27	0.015	381.3 ev	42.36	0.675	37.5	1.258×10^{15}	1.101×10^{15}	1.070×10^{16}	1.785×10^{16}
1.0	0.05	0.30	0.015	470.7 ev	47.07	0.607	30.4	1.918×10^{15}	1.386×10^{15}	1.495×10^{16}	2.247×10^{16}
1.2	0.10	0.36	0.030	678.0 ev	113	0.506	42.4	3.787×10^{15}	1.852×10^{15}	2.509×10^{16}	3.002×10^{16}
1.4	0.10	0.42	0.030	923.0 ev	132	0.434	31.0	6.194×10^{15}	2.041×10^{15}	3.552×10^{16}	3.309×10^{16}
1.6	0.10	0.48	0.030	1.21 kev	151	0.380	23.7	8.646×10^{15}	1.913×10^{15}	4.354×10^{16}	3.101×10^{16}
1.8	0.10	0.54	0.030	1.53 kev	170	0.337	18.7	1.047×10^{16}	1.554×10^{15}	4.697×10^{16}	2.519×10^{16}
2.0	0.10	0.60	0.030	1.88 kev	188	0.304	15.2	1.109×10^{16}	1.104×10^{15}	4.482×10^{16}	1.790×10^{16}
2.2	0.10	0.66	0.030	2.28 kev	207	0.276	12.5	1.051×10^{16}	7.015×10^{14}	3.861×10^{16}	1.137×10^{16}
2.4	0.10	0.72	0.030	2.71 kev	226	0.253	10.5	8.854×10^{15}	3.959×10^{14}	2.983×10^{16}	6.418×10^{15}
2.6	0.10	0.78	0.030	3.18 kev	245	0.233	8.98	6.665×10^{15}	1.999×10^{14}	2.072×10^{16}	3.240×10^{15}
2.8	0.10	0.84	0.030	3.69 kev	264	0.217	7.75	4.538×10^{15}	9.132×10^{13}	1.311×10^{16}	1.480×10^{15}
3.0	0.10	0.90	0.030	4.24 kev	282	0.203	6.75	2.799×10^{15}	3.757×10^{13}	7.541×10^{15}	6.090×10^{14}
3.2	0.10	0.96	0.030	4.82 kev	301	0.190	5.93	1.564×10^{15}	1.424×10^{13}	3.953×10^{15}	2.308×10^{14}
3.4	0.10	1.02	0.030	5.44 kev	320	0.179	5.26	7.932×10^{14}	5.033×10^{12}	1.887×10^{15}	8.159×10^{13}

TABLE 1

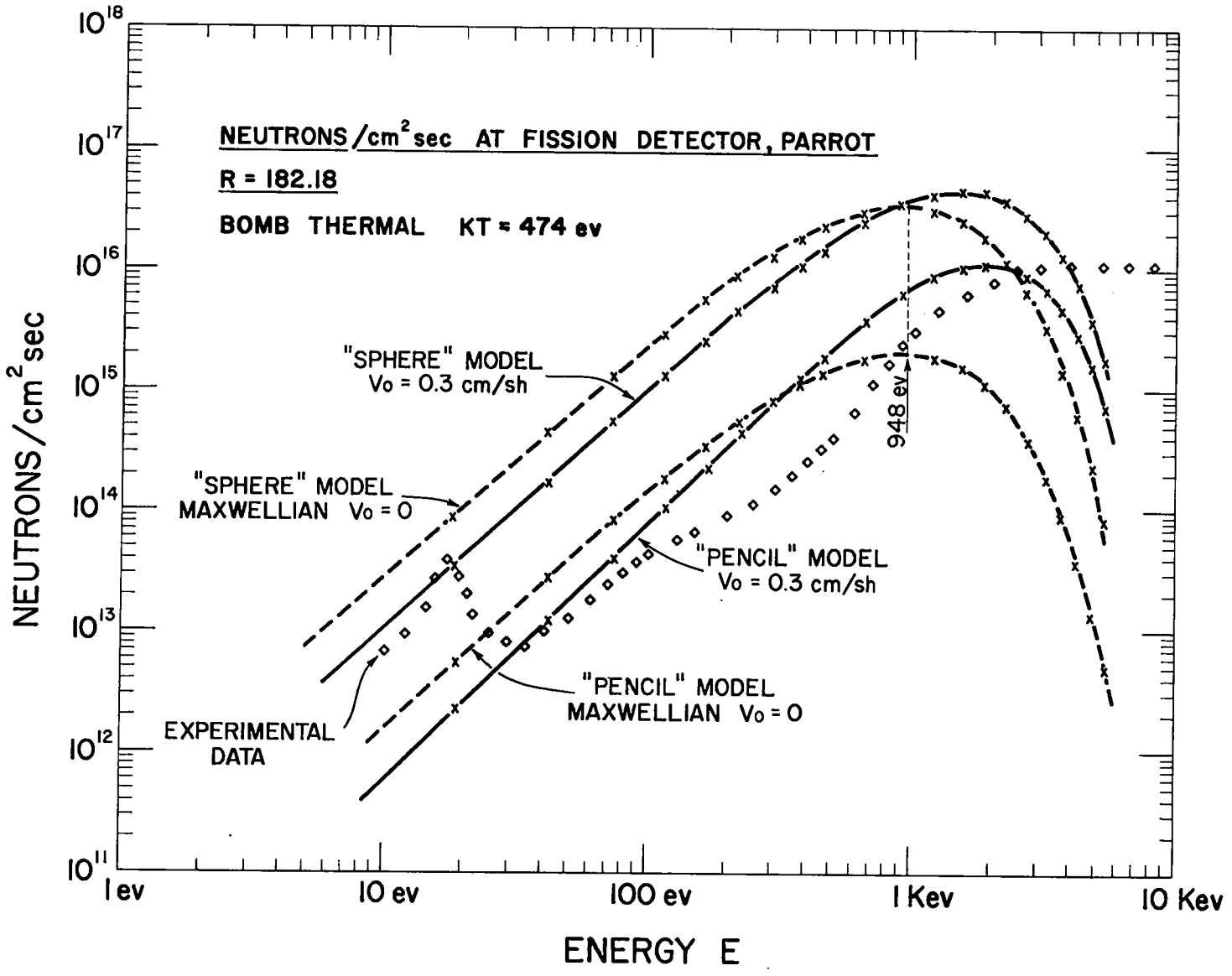

FIG. 6. THERMAL NEUTRON SPECTRA. BOMB THERMAL KT = 474 ev.

VI. EFFECT OF FINITE BOMB DEBRIS SPHERE ON MEASUREMENT

In our derivation of $\pi r^2 J_o(\vec{r},T)\Delta T$ for both the "pencil" and "sphere" model, Eqns. 14, 14a, 16, and 16a, we assume that the neutrons are released from a point source. Experimentally we count and determine the energy of the arriving neutrons solely by their arrival time.

Neutrons of energy E born either on the back or front side of the debris sphere with ΔR = 1 meter (Fig. 7) will therefore arrive at different times at the fission foil detector. This difference in time of flight will be interpreted wrongly as energy differences in the point model. We will discuss the expected experimental results for a neutron energy of 8.8 ev and assuming the debris dimensions shown in Fig. 7. This particular energy is chosen since U^{235} has a strong resonance at this energy.

Point Source, Case A.

We assume that the fission foil detector has a large and narrow resonance at E = 8.8 ev. All neutrons of energy E = 8.8 ev released from the point source (Fig. 7) will arrive at the fission foil detector at $T = 4.4416 \times 10^{-3}$ sec. Since this burst of neutrons has an energy of 8.8 ev, we will detect the neutrons properly by the U^{235} resonance at 8.8 ev in the fission foil.

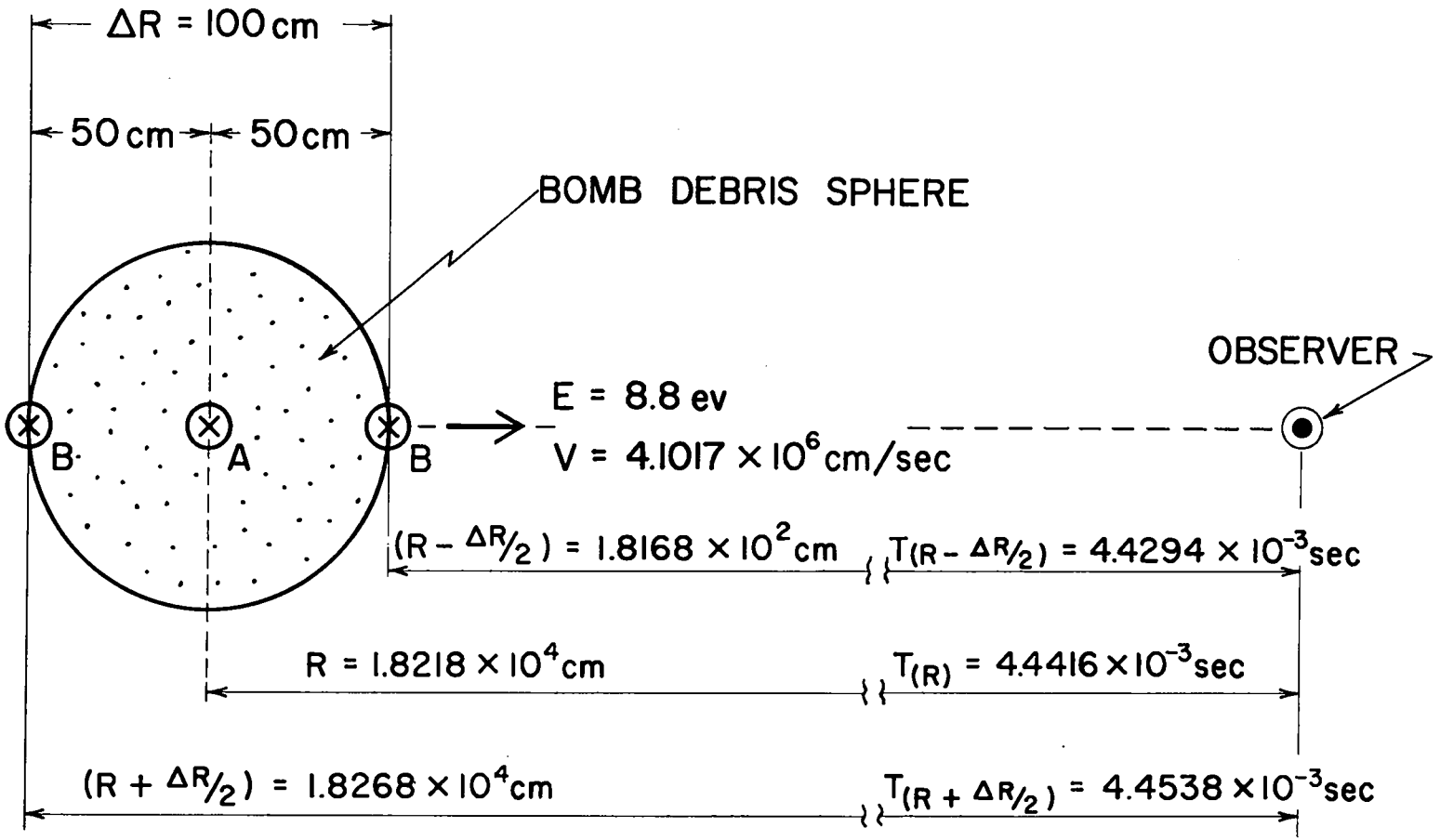

FIG. 7. MODEL A AND B FOR NEUTRON SOURCE IN DEBRIS SPHERE.

This correlation of the emitted neutron energy, the arrival time and the known resonance in the fission foil represents an effective checking mechanism on the neutron energy which arrives at the fission foil. Since the neutron arrival time is the "sole" quantity which determines the energy, we might easily be misled by neutrons of different energy which have been "bouncing" around in a non-direct pattern and just happen to arrive at the same time as the "direct" neutrons. This uncertainty can be discarded if one exposes a fission foil with known resonance structure. If one picks up these resonances, demonstrated by increased counts in the time sweep, we know that these are indeed neutrons corresponding to energies of the individual resonances.

Source at Front and Back, Case B.

Case A made the assumption that all neutrons are released from an ideal point source. We now assume that half the neutrons (E = 8.8 ev) of the point source are released from the front and the other half from the back of the bomb (Fig. 7). Since we determine neutron energy by arrival time, we will interpret the differences in arrival time due to different source locations as differences in the neutron energy. We ask what this apparent energy would be for both "front" and "back" neutrons if the neutrons have been released at the source with E = 8.8 ev. We also assume again that the fission foil has a resonance at this energy. Neutrons with E = 8.8 ev from the front arrive at 4.4294×10^{-3} sec and from the back at 4.4538×10^{-3} sec. The apparent energy difference is

$$\frac{\Delta E}{E} = \frac{2\Delta R}{R} = \frac{2x8.8x100}{1.8218x10^4} = 9.6608 \times 10^{-2} \qquad \text{EQN 21}$$

$$\Delta E_{(E=8.8\ ev)} = 0.0966\ ev.$$

Figure 8 shows the recorded counts for Case A and B.

If the separation of the two peaks (Case B) is larger than 0.08 ev, we can deduce that the source must be outside the debris sphere of $\Delta R = 100$ cm. This can be the case if fast neutrons escape the debris, and are subsequently reflected from the bottom of the hole which contained the device. These reflected neutrons can be moderated in the soil and will find their way back up the pipe. Cold neutrons can also pick up energy from the shock heated lining of the hole. Since the neutron energy is determined solely from arrival time, it will not tell us their past energy history. It is, therefore, quite difficult to accurately say much about neutrons which did not arrive at the detector in a straight time of flight fashion from the bomb debris.

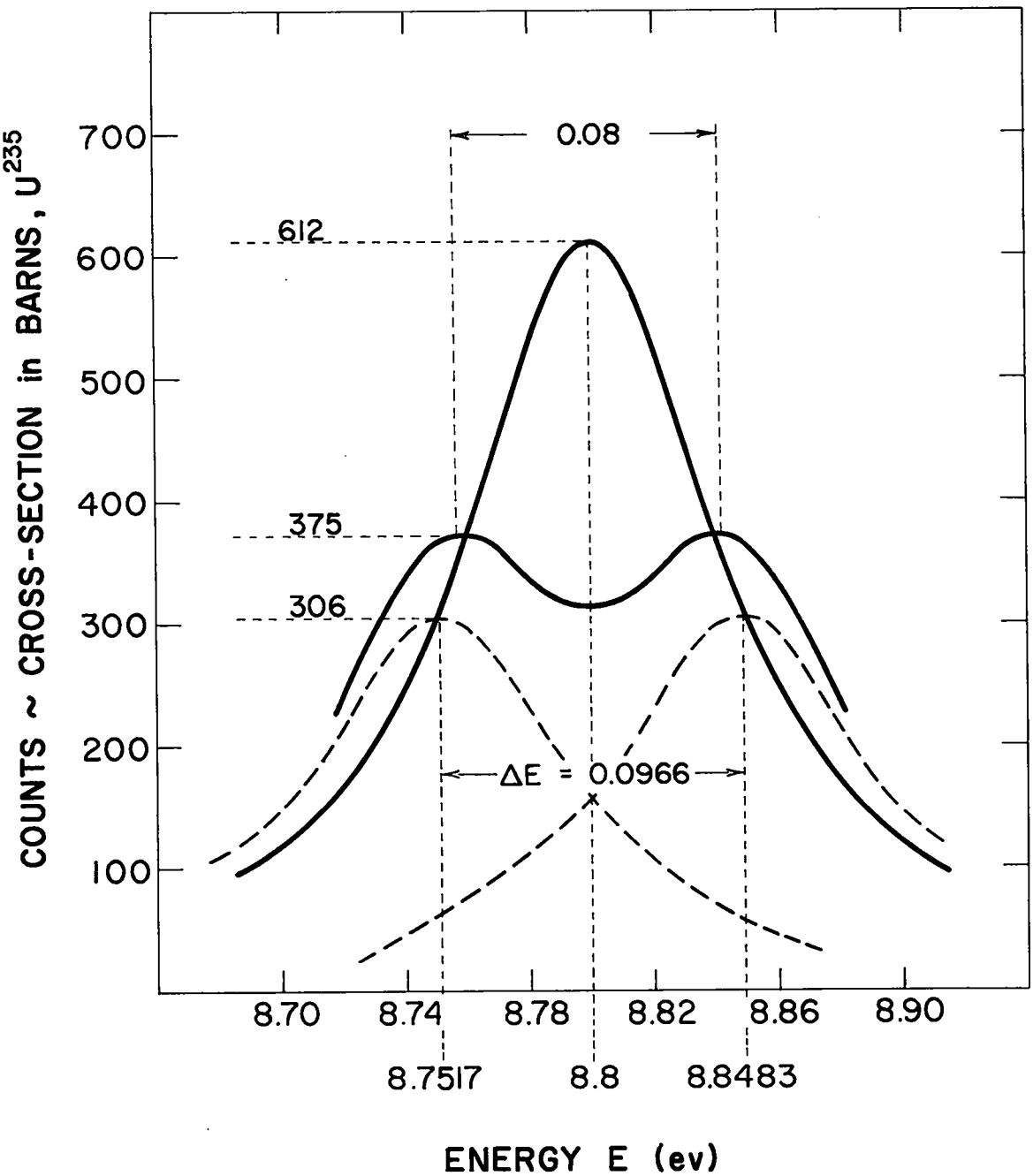

FIG. 8. COUNTS AT DETECTOR FOR DEBRIS SPHERE MODEL A AND B.

APPENDIX A

The distribution in flight time of neutrons reaching a detector remote from a pulsed source.

Theory.

We assume a localized neutron source $S(\vec{r},E,\vec{\Omega},t)$ according to Fig. A-1.

Boltzmann Equation.

$$\frac{1}{v}\frac{\partial\varphi(\vec{r},E,\vec{\Omega},t)}{\partial t}+\vec{\Omega}_o\cdot\nabla\varphi(\vec{r},E,\vec{\Omega},t)=S(\vec{r},E,\vec{\Omega},t) \qquad \text{EQN A-1}$$

Here the usual scattering term is missing since we assume vacuum between the source and the observer.

Integral Form.

$$\varphi(\vec{r},E,\vec{\Omega},t)=\int_{\ell} S\left(\vec{r}-\vec{\Omega}\ell,E,\vec{\Omega},t-\frac{\ell}{v}\right)d\ell \qquad \text{EQN A-2}$$

Here neutrons which leave the source at $t - \ell/v$ with constant velocity v arrive at the observer at time t. In the integral form the second term in Eqn. A-1 becomes zero due to the assumed δ-function source pulse width.

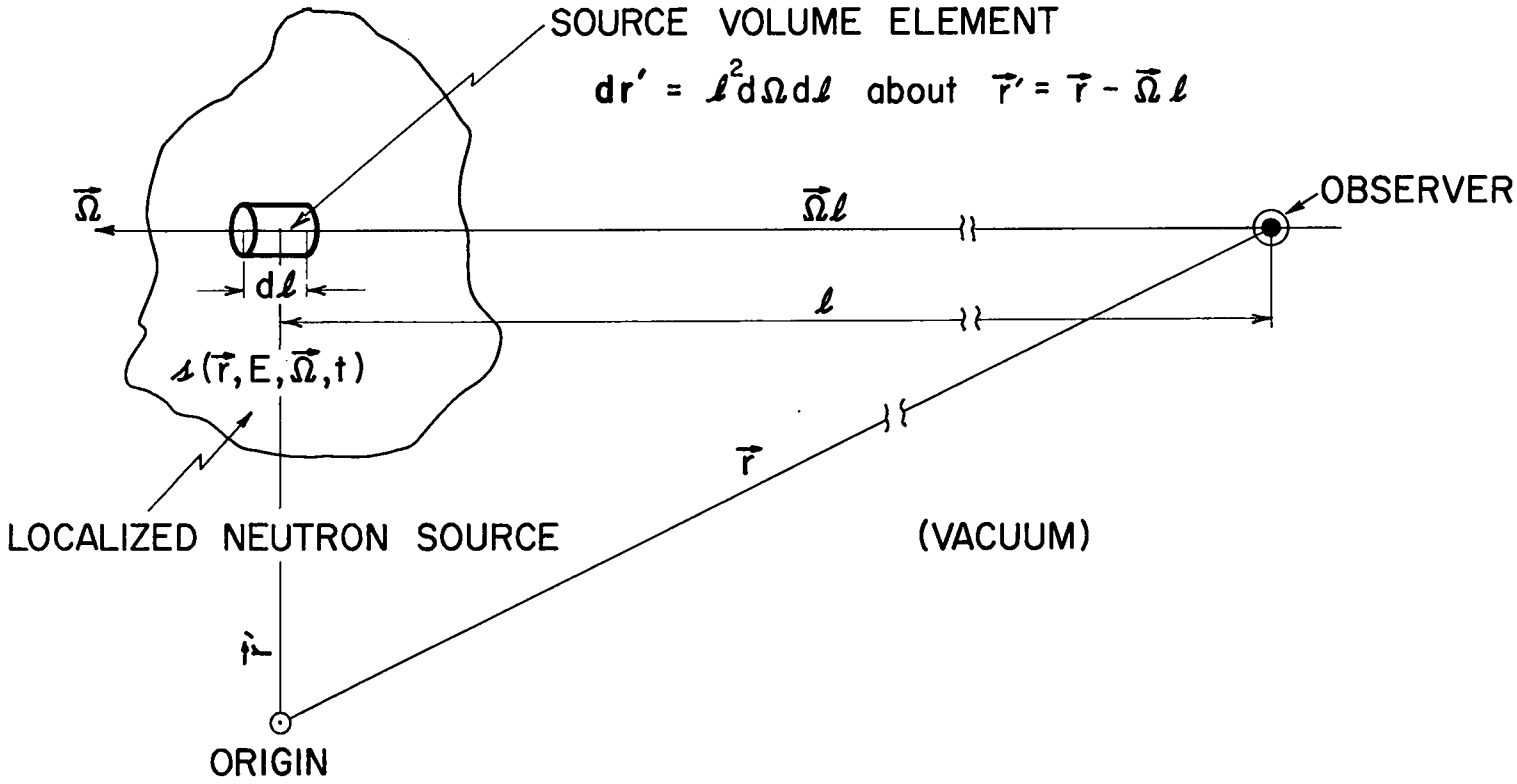

FIG. A-1. SOURCE--OBSERVER CONFIGURATION.

Total flux at point of observer

$$\Phi(\vec{r},E,t) = \int_{\vec{\Omega}} \varphi(\vec{r},E,\vec{\Omega},t)d\Omega = \int_{\ell}\int_{\vec{\Omega}} S\left(\vec{r}-\vec{\Omega}\ell,E,\vec{\Omega},t-\frac{\ell}{v}\right)d\ell d\Omega \qquad \text{EQN A-3}$$

$$= \int_{\text{source}} S\left(\vec{r}',E,\frac{\vec{r}-\vec{r}'}{\ell},t-\frac{\ell}{v}\right)\frac{dr'}{\ell^2}$$

Total current at point of observer

$$\vec{J}(\vec{r},E,t) = \int_{\vec{\Omega}} \varphi(\vec{r},E,\vec{\Omega},t)\vec{\Omega}d\Omega$$

$$= \int_{\ell}\int_{\vec{\Omega}} S\left(\vec{r}-\vec{\Omega}\ell,E,\vec{\Omega},t-\frac{\ell}{v}\right)\vec{\Omega}d\ell d\Omega$$

$$= \int_{\text{source}} \left(\frac{\vec{r}-\vec{r}'}{\ell}\right) S\left(\vec{r}',E,\frac{\vec{r}-\vec{r}'}{\ell},t-\frac{\ell}{v}\right)\frac{dr'}{\ell^2} \qquad \text{EQN A-4}$$

Delta Function Source.

$$S(\vec{r}',E,\vec{\Omega},t) = S_o(\vec{r}',E,\vec{\Omega})\delta(t) \qquad \text{EQN A-5}$$

We now integrate the Boltzmann equation, Eqn. A-1, from $0^- \leq t \leq 0^+$ and obtain the initial neutron population created by the δ-function source at $\vec{r}'$.

$$\frac{\varphi(\vec{r}',E,\vec{\Omega},0^+)}{v} = n(\vec{r}',E,\vec{\Omega},0^+) = S_o(\vec{r}',E,\vec{\Omega}) \qquad \text{EQN A-6}$$

From Eqn. A-6 and Eqn. A-3 we set for the total flux at the point of observer $\vec{r}$.

$$\Phi(\vec{r},E,t) = \int_{\text{source}} n\left(\vec{r}',E,\frac{\vec{r}-\vec{r}'}{\ell},0^+\right)\delta\left(t-\frac{\ell}{v}\right)\frac{dr'}{\ell^2} \qquad \text{EQN A-7}$$

The total flux Φ at the observer is due to the initial neutron density n at time 0^+ on the spherical surface $(\vec{r}'-\vec{r}) = \ell$.

For a localized source and distant field point $\vec{r}$, say the distance R from the source center, what is the total current passing through a unit area at $\vec{r}$ in the time interval ΔT at T? The above statement implies that the source dimension ΔR is small compared with R, the distance between the source and the observer. ΔT is, therefore, large compared to the neutron transit time across the source (assuming constant velocity) but still small (increment) compared to the flight time from the source to the observer.

$$\vec{J}_o(\vec{r},T)\Delta T = \int_{T-\frac{\Delta T}{2}}^{T+\frac{\Delta T}{2}} dt \int_o^\infty dE\ \vec{J}(\vec{r},E,t)$$

$$= \int_{T-\frac{\Delta T}{2}}^{T+\frac{\Delta T}{2}} dt \int_o^\infty dE \int_{\text{source}} \left(\frac{r-r'}{\ell}\right) n\left(\vec{r}',E,\frac{\vec{r}-\vec{r}'}{\ell},0^+\right)\delta\left(t-\frac{\ell}{v}\right)\frac{dr'}{\ell^2}$$

EQN A-8

If one integrates Eqn. A-8 over dt first, we get

$$\vec{J}_o(\vec{r},T)\Delta T = \int_o^\infty dE \int_{\text{source}} \left(\frac{\vec{r}-\vec{r}'}{\ell}\right) n\left(\vec{r}',E,\frac{\vec{r}-\vec{r}'}{\ell},0^+\right)\frac{dr'}{\ell^2} \quad \text{EQN A-9}$$

where the range of E or v and $\ell(r')$ or R is restricted by

$$T - \frac{\Delta T}{2} \leq \frac{\ell(\vec{r}')}{v} \leq T + \frac{\Delta T}{2} \quad \text{EQN A-10}$$

Equation A-10 expresses the time of flight concept; that is, the energy of neutrons arriving at the observer is solely determined by the neutron arrival time T.

The neutron energy or velocity which is deduced from the flight time T is

$$v_{max} = \frac{R+\Delta R/2}{T-\Delta T/2} \simeq \frac{R}{T-\Delta T/2}$$

$$v_{min} = \frac{R-\Delta R/2}{T+\Delta T/2} \simeq \frac{R}{T+\Delta T/2}$$

or $$\Delta v = \frac{R}{T^2}\Delta T \quad \text{EQN A-11}$$

To better see the meaning of this approximation, we sketch Fig. A-2 on the ℓ,v plane, that region containing $t - \ell/v = 0$ for $T - \frac{\Delta T}{2} \leq t \leq T + \frac{\Delta T}{2}$ and $R - \frac{\Delta R}{2} \leq \ell \leq R + \frac{\Delta R}{2}$.

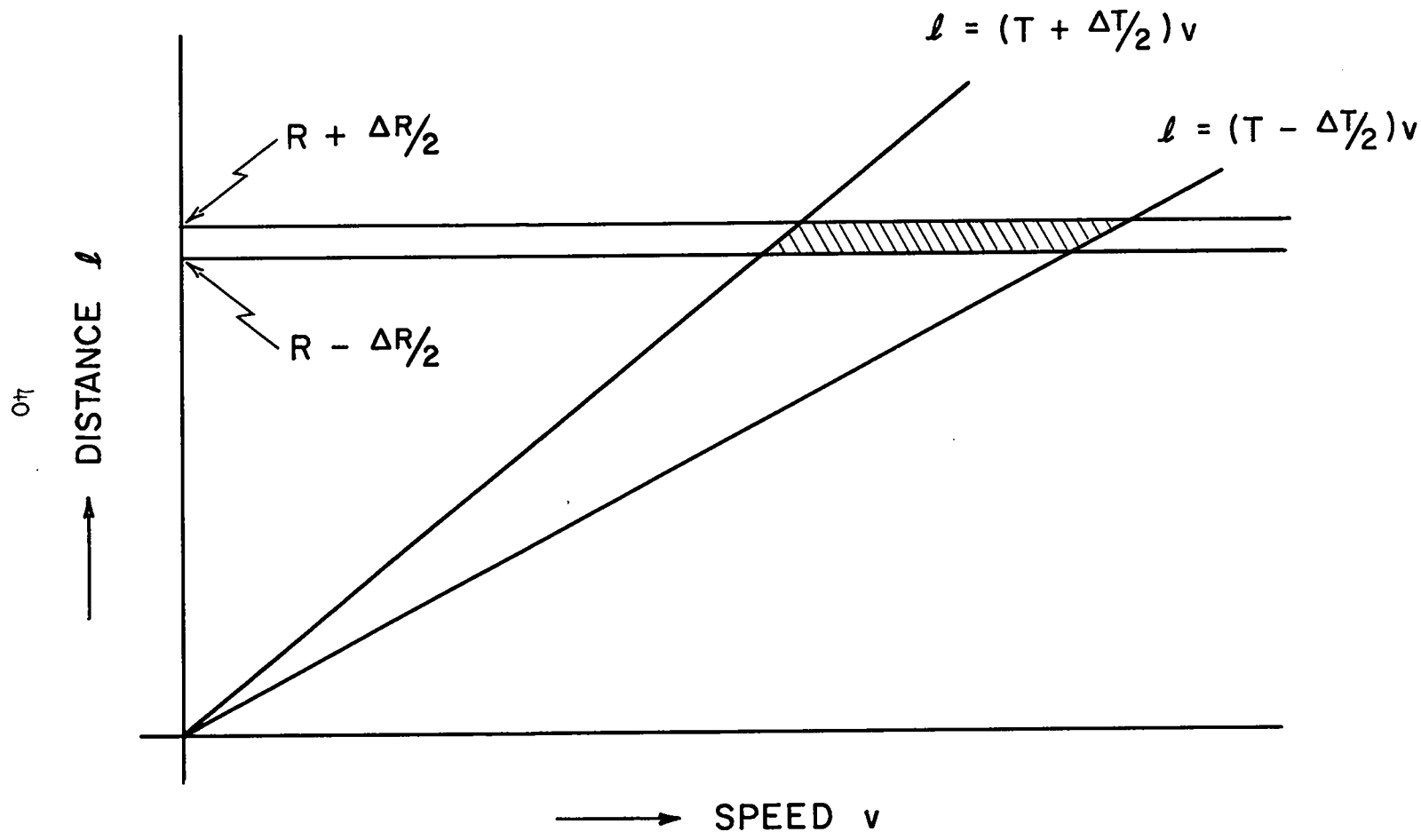

FIG. A-2. FINITE SOURCE ΔR AND TIME OF FLIGHT.

The hashed area element in Fig. A-2 indicates the loci of ℓ,v points which permit neutrons to reach the observer at t = T in $\pm$ $\Delta T/2$. Our above assertion that ΔT be large compared to the transit time of a neutron across the source (i.e. $\Delta T \gg \Delta R/v$) implies that the change in v along the $\ell = (T \pm \Delta T/2)v$ borders of this area element is small compared to the change $\Delta v \cong v_{max} - v_{min}$, along the $\ell = R \pm \Delta R/2$ borders. The approximation, that v and T (or E and T) are functionally related, breaks down for the time resolution $\Delta T = \Delta R/v$, and this corresponds to a velocity (speed) or energy resolution

$$\frac{\Delta v}{v} \cong \frac{\Delta R}{R} \cong \frac{\Delta E}{2E} \qquad \text{EQN A-12}$$

For the case of interest in this report $\frac{\Delta R}{R} \simeq 0.005$. For E = 10 ev we get therefore for $\Delta E = 0.108$ ev. For the above numerical example the uncertainty in the source position produces an uncertainty in the energy of 0.1 ev.

In Eqn. A-9 we change the variable E to v.

$$\vec{J}_o(\vec{r},T)\Delta T = \int_{source} dr' \left(\frac{\vec{r}-\vec{r}'}{\ell}\right) n\, \vec{r}', v, \frac{\vec{r}-\vec{r}'}{\ell}, 0^+ \frac{\Delta v}{\ell^2} \qquad \text{EQN A-13}$$

Since

$$\left(\frac{\vec{r}-\vec{r}'}{\ell}\right) \simeq \frac{(\vec{r}-\vec{r}'_{ct.})}{R} = \vec{\Omega}_o$$

and

$$\vec{J}_o = J_o \vec{\Omega}_o$$

$$J_o(\vec{r},T)\Delta T = \int_{\text{source}} dr' \, n(\vec{r}',v,\vec{\Omega}_o,o^+) \frac{\Delta v}{R^2} \qquad \text{EQN A-14}$$

where $$v = \frac{R}{T} = \sqrt{\frac{2E}{m}}$$

$$\Delta v = R \frac{\Delta T}{T^2} \; ; \quad \Delta E = \frac{mR}{T} \Delta v$$

Equation A-14 gives the number of neutrons passing through the unit area perpendicular to $\vec{\Omega}_o$ at $\vec{r}$ in the time interval ΔT about T in terms of that portion of the initial neutron population at $\vec{r}'$ with energies specified by Δv about v. The conversion from the time distribution function, $J_o(\vec{r},T)$, to an energy distribution function, say $J_o(\vec{r},E)$ neutrons per cm^2 per unit energy interval, or a velocity (speed) distribution function, say $J_o(\vec{r},v)$, is readily accomplished through the obvious relations

$$J_o(\vec{r},T)\Delta T = J_o(\vec{r},E)\Delta E = J_o(\vec{r},v)\Delta v \qquad \text{EQN A-15}$$

APPENDIX B

Total source strength.

We want to determine factor c in Eqn. 1. The total neutrons produced are

$$N_o = \int_{\text{source}} dr' \int_v dv \int_{\vec{\Omega}} d\Omega \int_t dt\, S(r', v, \vec{\Omega}, t) \qquad \text{EQN B-1}$$

The initial neutrons are produced instantaneously by a δ-function source

$$N_o = \int_{\text{source}} dr' \int_v dv \int_{\vec{\Omega}} d\Omega\, n(\vec{r}', v, \vec{\Omega}, o^+) \qquad \text{EQN B-2}$$

The initial neutron population $n(\vec{r}', v, \vec{\Omega}, o^+)$ created by the δ-function source is as in Eqn. 2.

$$n(\vec{r}', v, \vec{\Omega}, o^+) = ce^{-\frac{m}{2KT}(v^2+v_o^2-2vv_o\cos\psi_o)} \times v^2 \qquad \text{EQN B-3}$$

In Eqn. 2 we had the preferred direction $\vec{\Omega}_o$, but here we want to integrate over all directions $\vec{\Omega}$. We also make the same assumption as in Eqn. 3 that the neutrons are produced uniformly over the sphere.

Inserting Eqn. B-3 into Eqn. B-2 we get

$$N_o = \int_{\text{source}} dr' \int_v dv \int_{-1.0}^{+1.0} 2\pi d\cos\psi_o \; ce^{-\frac{m}{2KT}(v^2+v_o^2-2vv_o\cos\psi_o)} \times v^2 \qquad \text{EQN B-4}$$

Here the limits of the last integral are inverted since $d\Omega = 2\pi \sin\psi_o d\psi_o = -2\pi d\cos\psi_o$. The last integral over the angle ψ_o we can readily integrate over its limits. We get

$$N_o = \int_{\text{source}} dr' \int_o^\infty dv \, \frac{2\pi cvKT}{mv_o} \left[e^{-\frac{m}{2KT}(v-v_o)^2} - e^{-\frac{m}{2KT}(v+v_o)^2} \right] \qquad \text{EQN B-5}$$

First we integrate Eqn. B-2 over v. We integrate the two exponential terms separately. The first term is

$$\int_o^\infty vdve^{-\frac{m}{2KT}(v-v_o)^2} = \int_o^\infty (v-v_o)d(v-v_o)e^{-\frac{m}{2KT}(v-v_o)^2}$$

$$+ \int_o^\infty v_o d(v-v_o)e^{-\frac{m}{2KT}(v-v_o)^2}$$

$$= \frac{KT}{m}\int_{x_o}^\infty dx_1 e^{-x_1} + v_o\sqrt{\frac{2KT}{m}}\int_{-y_o}^\infty dy_1 e^{-y_1^2} \qquad \text{EQN B-6}$$

The second term is

$$\int_o^\infty vdve^{-\frac{m}{2KT}(v+v_o)^2} = \int_o^\infty (v+v_o)d(v+v_o)e^{-\frac{m}{2KT}(v+v_o)^2}$$

$$- \int_o^\infty v_o d(v+v_o)e^{-\frac{m}{2KT}(v+v_o)^2}$$

$$= \frac{KT}{m}\int_{x_o}^\infty dx_2 e^{-x_2^2} - v_o\sqrt{\frac{2KT}{m}}\int_{y_o}^\infty dy_2 e^{-y_2^2} \qquad \text{EQN B-7}$$

The difference between Eqn. B-6 and Eqn. B-7 is

$$v_o\sqrt{\frac{2KT}{m}}\int_{-\infty}^{+\infty} dye^{-y^2} = v_o\sqrt{\frac{2\pi KT}{m}} \qquad \text{EQN B-8}$$

Equation B-5 now becomes

$$N_o = \frac{2\pi cKT}{m}\sqrt{\frac{2\pi KT}{m}}\int_{\text{source}} dr' \qquad \text{EQN B-9}$$

We can now determine the constant c

$$c = \frac{N_o}{V_o}\left(\frac{2\pi KT}{m}\right)^{-3/2} \qquad \text{EQN B-10}$$

where N_o = Total thermal neutrons produced in sphere

V_o = Source volume in cm^3

$\sqrt{\frac{KT}{m}}$ = Most probable neutron velocity in Boltzmann distribution

APPENDIX C

Relationship between N(t) (neuts/cm^2/sec) at fission foil and n(E) (neuts/Mev) at point of explosion.

The following relationship holds

$$N(t)dt = n(E)dE \frac{\pi r^2}{4\pi R^2} \qquad \text{EQN C-1}$$

where $\pi r^2/4\pi R^2 = 2.486 \times 10^{-10}$ geometric attenuation factor. We first determine dt

$$t_{(sec)} = \frac{R}{v} = \frac{1.8218 \times 10^4 \text{ cm}}{13.827 \sqrt{E(\text{Mev})} \times 10^8 \text{ (cm/sec)}}$$

$$= 1.317 \times 10^{-5} \, E(\text{Mev})^{-1/2}$$

$$dt = E^{-3/2} \, dE \, \frac{1.317 \times 10^{-5}}{2}$$

Substituting dt into Eqn. 26

$$N(t) = n(E) \frac{dE}{dt} = n(E) \, E^{3/2} \, \frac{2}{1.317 \times 10^{-5}} \times 2.486 \times 10^{-10}$$

We finally get

$$N(t) = 3.775 \times 10^{-5} E^{3/2}(\text{Mev})\, n(E) \qquad \text{EQN C-2}$$

Equation C-2 is the conversion formula to convert virgin escape neutrons from the bomb $n(E)$ in neuts/Mev to $N(t)$, the neutron flux in neutrons/cm^2 sec at the fission foil of area $\pi r^2 = 1.037\ cm^2$ at $R = 182.18$ m. The quantity $n(E)$ in neuts/Mev is a convenient quantity in bomb diagnostics, since the integral over energy of $n(E)$ represents the total number of neutrons leaking from the device, which can be directly related to such quantities as the yield of Parrot and fission neutrons born.